

Figure 9. from Mapping Distances across the Perseus Molecular Cloud Using CO Observations, Stellar Photometry, and Gaia DR2 Parallax Measurements

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$d_{fore}$	$E_{fore}$	$P_b$	$d_1$	$d_2$	$d_3$	$d_4$	$d_5$	$c_1$	$c_2$	$c_3$	$c_4$	$c_5$
$5.64^{+0.25}_{-0.56}$	$0.16^{+0.00}_{-0.01}$	$0.02^{+0.00}_{-0.00}$	$7.30^{+0.11}_{-0.12}$	$7.30^{+0.08}_{-0.08}$	$7.34^{+0.08}_{-0.08}$	$7.28^{+0.07}_{-0.06}$	$7.34^{+0.07}_{-0.08}$	$0.90^{+0.07}_{-0.07}$	$0.36^{+0.02}_{-0.02}$	$0.37^{+0.03}_{-0.03}$	$0.91^{+0.02}_{-0.02}$	$1.15^{+0.05}_{-0.04}$
			$-0.7 \text{ km s}^{-1}$	$0.8 \text{ km s}^{-1}$	$2.9 \text{ km s}^{-1}$	$4.8 \text{ km s}^{-1}$	$6.6 \text{ km s}^{-1}$					

