

Figure 8. from Mapping Distances across the Perseus Molecular Cloud Using CO Observations, Stellar Photometry, and Gaia DR2 Parallax Measurements

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$d_{fore}$	$E_{fore}$	$P_b$	$d_1$	$d_2$	$d_3$	$d_4$	$d_5$	$d_6$	$c_1$	$c_2$	$c_3$	$c_4$	$c_5$	$c_6$
$5.20^{+0.43}_{-0.85}$	$0.10^{+0.01}_{-0.01}$	$0.02^{+0.00}_{-0.00}$	$7.08^{+0.70}_{-0.34}$	$7.40^{+0.04}_{-0.04}$	$7.29^{+0.15}_{-0.10}$	$7.37^{+0.06}_{-0.10}$	$7.39^{+0.03}_{-0.02}$	$7.42^{+0.05}_{-0.05}$	$0.33^{+0.03}_{-0.04}$	$0.97^{+0.03}_{-0.03}$	$0.77^{+0.03}_{-0.03}$	$0.44^{+0.02}_{-0.02}$	$0.83^{+0.01}_{-0.01}$	$0.32^{+0.02}_{-0.02}$
			$-0.1 \text{ km s}^{-1}$	$1.7 \text{ km s}^{-1}$	$3.5 \text{ km s}^{-1}$	$5.5 \text{ km s}^{-1}$	$7.3 \text{ km s}^{-1}$	$8.8 \text{ km s}^{-1}$						

Distance (pc)

