

Figure 5. from Mapping Distances across the Perseus Molecular Cloud Using CO Observations, Stellar Photometry, and Gaia DR2 Parallax Measurements

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$d_{\text{fore}}$	$E_{\text{fore}}$	$P_b$	$d_1$	$d_2$	$d_3$	$c_1$	$c_2$	$c_3$
$4.50^{+0.39}_{-0.32}$	$0.25^{+0.01}_{-0.01}$	$0.02^{+0.00}_{-0.00}$	$7.56^{+0.12}_{-0.11}$	$7.40^{+0.05}_{-0.06}$	$7.34^{+0.10}_{-0.08}$	$1.34^{+0.08}_{-0.08}$	$1.24^{+0.02}_{-0.03}$	$0.47^{+0.03}_{-0.03}$
			$8.3 \text{ km s}^{-1}$	$9.7 \text{ km s}^{-1}$	$11.0 \text{ km s}^{-1}$			

