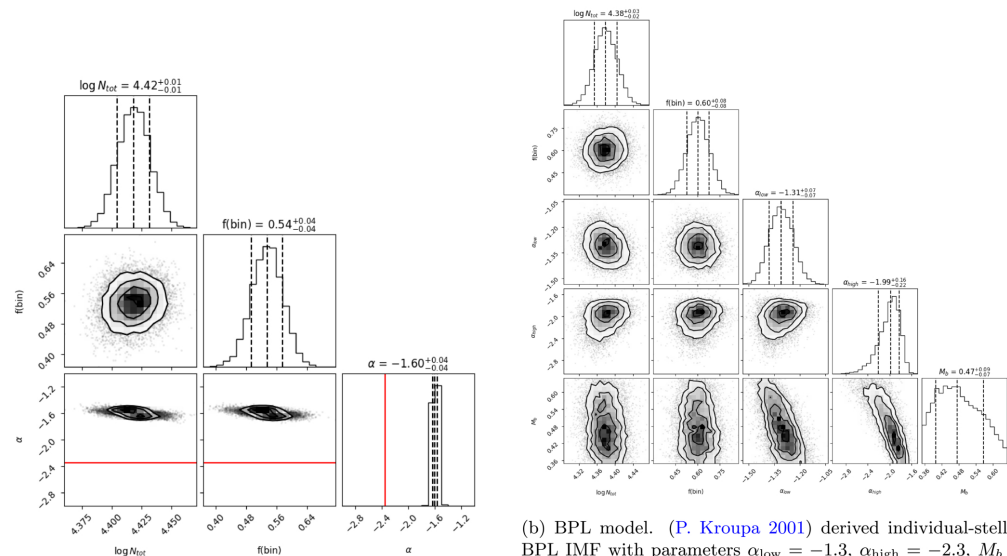
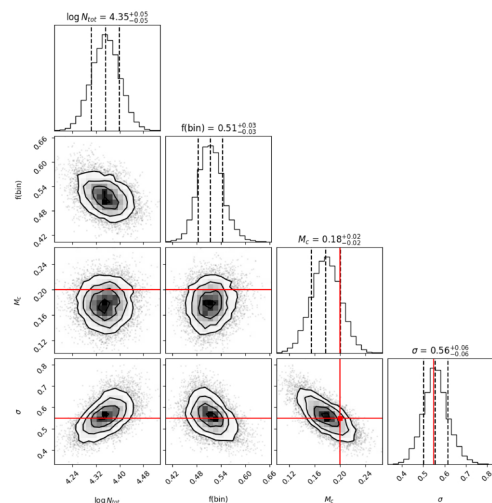


Figure 6. from
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(a) SPL model. The reference slope $\alpha = -2.35$ (E. E. Salpeter 1955) is shown in red.

(b) BPL model. (P. Kroupa 2001) derived individual-stellar BPL IMF with parameters $\alpha_{\text{low}} = -1.3$, $\alpha_{\text{high}} = -2.3$, $M_b = 0.5M_{\odot}$. Since our analysis measures the system IMF without resolving or correcting for stellar multiplicity, these parameters are not directly comparable.



(c) LN model. Reference values $M_c = 0.2M_{\odot}$, $\sigma = 0.55$ (G. Chabrier 2003) are shown in red.