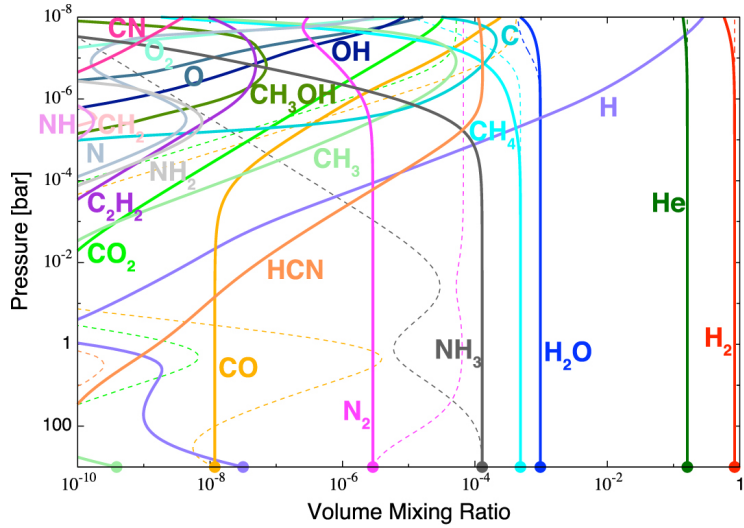


Figure 16. from Theoretical Transmission Spectra of Exoplanet Atmospheres with Hydrocarbon Haze: Effect of Creation, Growth, and Settling of Haze Particles. II. Dependence on UV Irradiation Intensity, Metallicity, C/O Ratio, Eddy Diffusion Coefficient, and Temperature

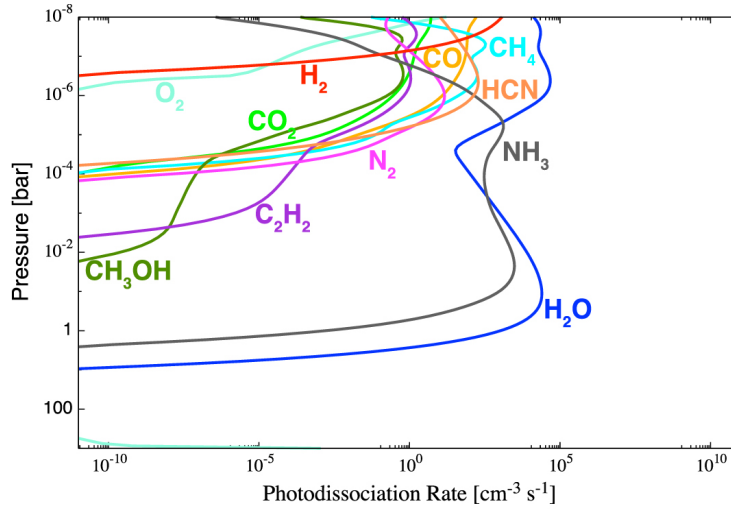
null 2019 APJ 877 109 doi:10.3847/1538-4357/ab1b1d

<https://dx.doi.org/10.3847/1538-4357/ab1b1d>

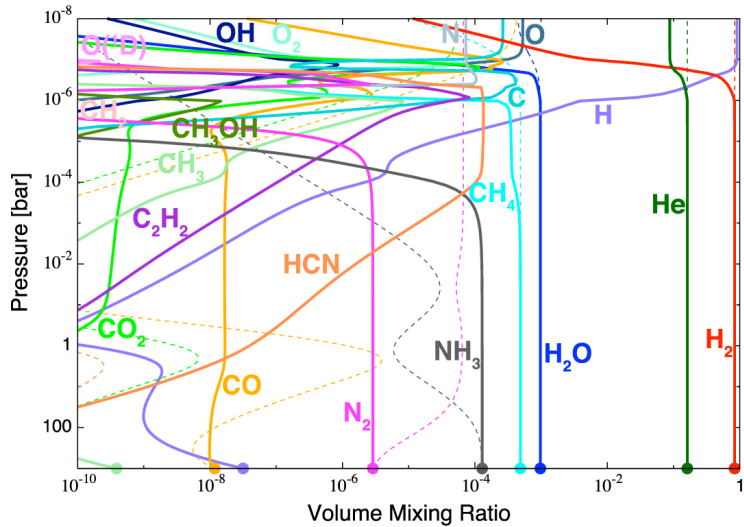
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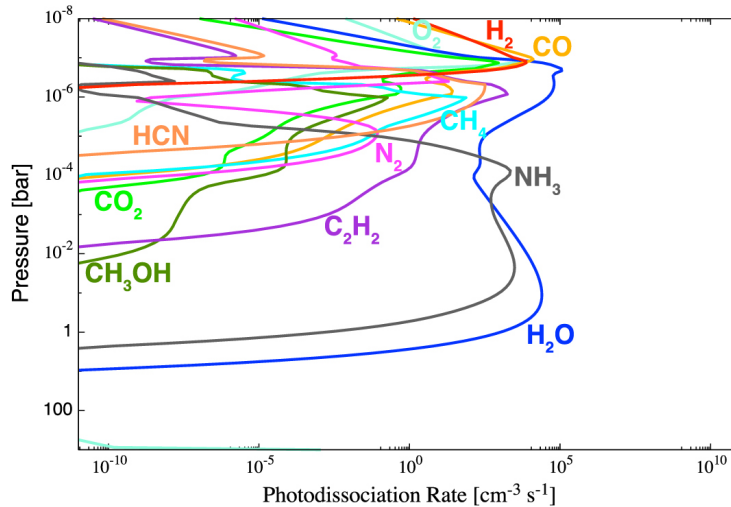
(a) $K_{zz} = 1 \times 10^9 \text{ cm}^2 \text{ s}^{-1}$



(b) $K_{zz} = 1 \times 10^9 \text{ cm}^2 \text{ s}^{-1}$



(c) $K_{zz} = 1 \times 10^5 \text{ cm}^2 \text{ s}^{-1}$



(d) $K_{zz} = 1 \times 10^5 \text{ cm}^2 \text{ s}^{-1}$