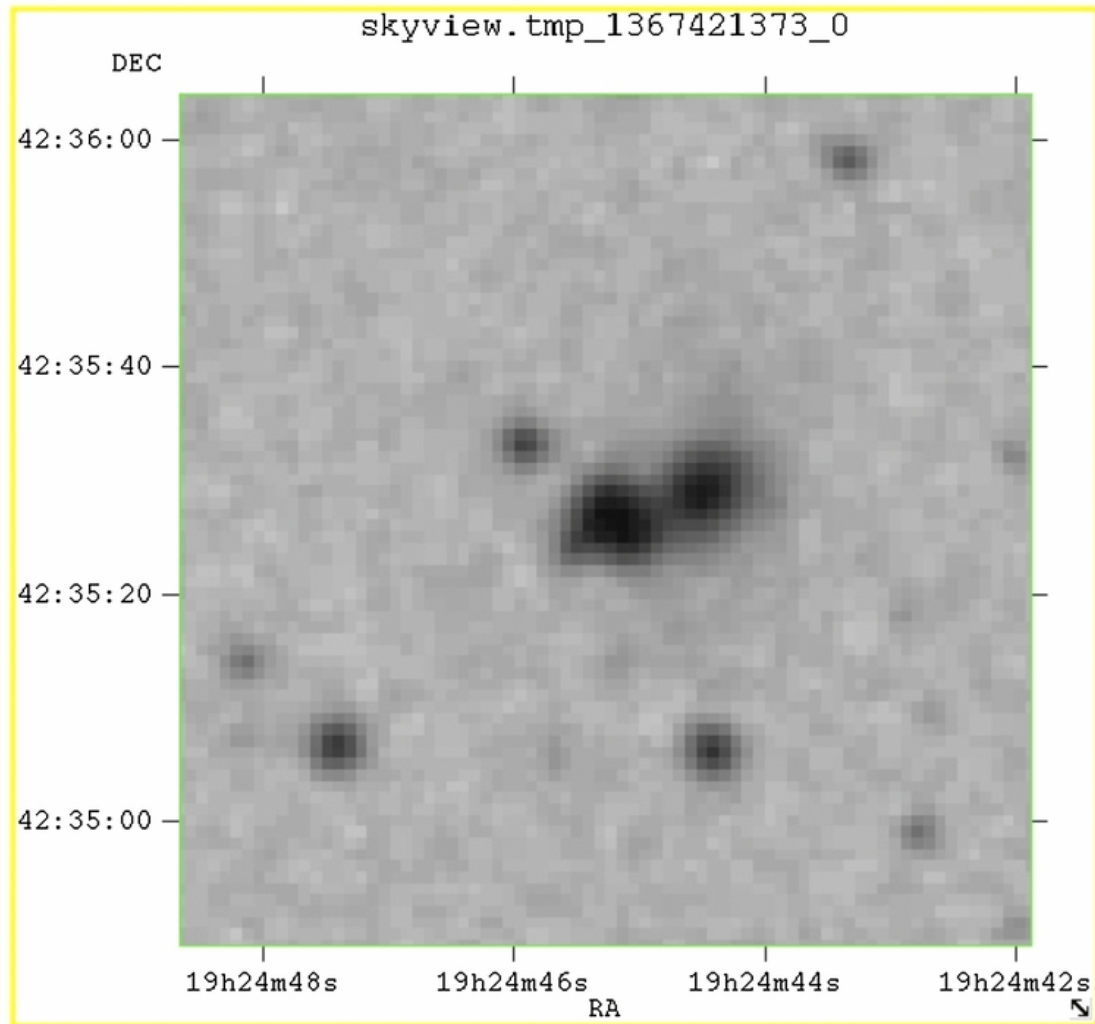


Figure 3. from A Technique to Derive Improved Proper Motions for Kepler Objects of Interest

Benedict et al. 2014 AJ 148 108 doi:10.1088/0004-6256/148/6/108

<https://dx.doi.org/10.1088/0004-6256/148/6/108>

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