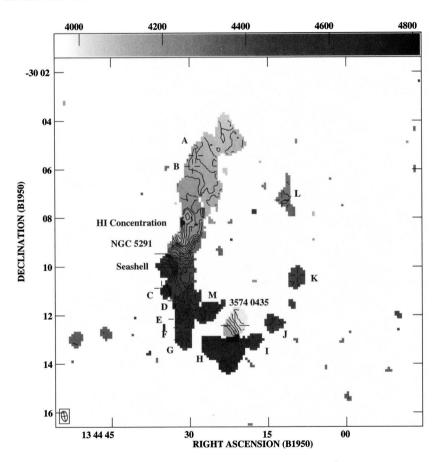
Figure from NGC 5291: Implications for the Formation of Dwarf Galaxies Malphrus 1997 AJ 114 1438 doi:308158864287067800170087\_doi:10.1086/118574 https://dx.doi.org/308158864287067800170087\_doi:10.1086/118574 The Astronomical Journal

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noment) map from the C+D combined array data. Isovelocity contours are 25 km s<sup>-1</sup> apart and run from 4000 km s<sup>-</sup>

eviously too-tenuous H I disk. that the H I offset, at least, now appears to d distribution of the gas as seen at low nematics, but especially the morphology of stem as revealed by the higher resolution ed here, encourage us to reinstate as a posbased on analysis of the existing optical, raidata

4.1.1 Ram sweeping

As NGC 5291 is located in the northwest pa