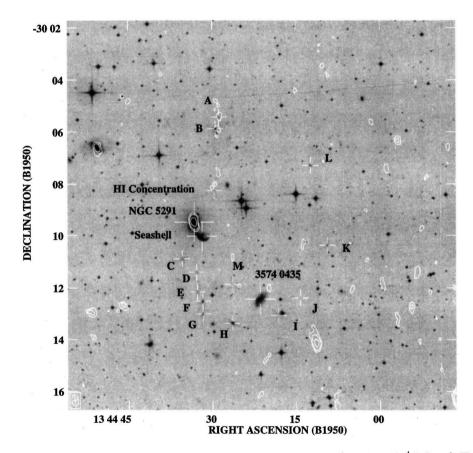
Figure from NGC 5291: Implications for the Formation of Dwarf Galaxies Malphrus 1997 AJ 114 1438 doi:308158864287067800170087_doi:10.1086/118574 https://dx.doi.org/308158864287067800170087_doi:10.1086/118574 The Astronomical Journal

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n contours over optical image. Contour levels correspond to 3, 5, and 10σ ($1\sigma=2.7\times10^{-4}$ Jy/beam). The beams in the lower right corner. Optical and H I knots are indicated with labeled crosses.

ghtness temperature in Kelvin and v is the The peak column densities found for the (the galaxies and knots) on the moment Table 3. These values were determined as data which is most sensitive to ex-

The integrated flux from the H I gas may integrating Eq. (1) over the solid angle subtreature, and expressing the brightness tempers of the observed flux S(v). The H I mass may treated as follows (Mihalas & Binney 1981):