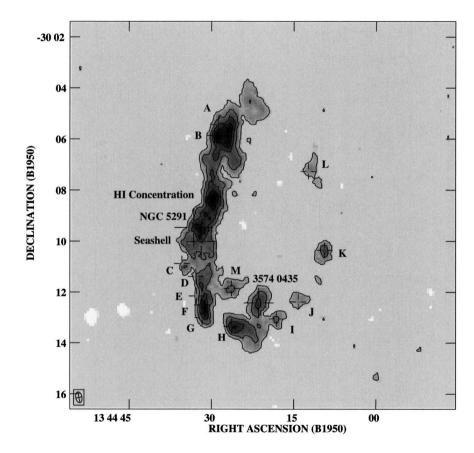
Figure from NGC 5291: Implications for the Formation of Dwarf Galaxies Malphrus 1997 AJ 114 1438 doi:308158864287067800170087_doi:10.1086/118574 https://dx.doi.org/308158864287067800170087_doi:10.1086/118574 The Astronomical Journal

S ET AL.: NGC 5291



e of the H I integrated intensity from the C+D combined array data. Contour levels are $1.1 (2\sigma)$, 5 10, 15, and $20 \times$ ") is indicated by the ellipse in the lower right corner. Optical and H I knots are indicated with labeled crosses.

associated with optical knots. sciation of the H I concentrations with the s is very striking. Figure 3 shows that the the complex is generally clumpy; the

troscopy is needed to verify these associations A strong, highly resolved H I feature (cleaving. 3) lies \sim 0.5 west of a point about a thin from the nucleus of NGC 5291 to the brightest