

Figure 10. from Mapping Distances across the Perseus Molecular Cloud Using CO Observations, Stellar Photometry, and Gaia DR2 Parallax Measurements

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$d_{fore}$	$E_{fore}$	$P_b$	$d_1$	$d_2$	$d_3$	$d_4$	$c_1$	$c_2$	$c_3$	$c_4$
$4.09^{+0.10}_{-0.06}$	$0.07^{+0.01}_{-0.01}$	$0.02^{+0.00}_{-0.00}$	$7.24^{+0.15}_{-0.17}$	$6.74^{+0.10}_{-0.11}$	$7.33^{+0.04}_{-0.03}$	$6.72^{+0.20}_{-0.15}$	$0.53^{+0.03}_{-0.04}$	$0.32^{+0.02}_{-0.02}$	$1.30^{+0.04}_{-0.03}$	$0.30^{+0.03}_{-0.03}$
			$0.3 \text{ km s}^{-1}$	$2.3 \text{ km s}^{-1}$	$4.2 \text{ km s}^{-1}$	$5.8 \text{ km s}^{-1}$				

